

# On Static Charged Fluid Sphere in General Relativity

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## ABSTRACT

The present paper provides solution of Einstein-Maxwell field equations for static charged fluid sphere using specific choice of  $g_{11}$  and  $g_{44}$ .

## 1. Introduction

Shi chang [6] has found some conformal flat interior solutions of the Einstein- Maxwell equations for a charged stable static sphere. These solutions satisfy physical conditions inside the sphere. Some exact static solutions of Einstein- Maxwell equations representing a charged fluid sphere were obtained by Singh & Yadav[5]. Xingxiang[7] has obtained an exact solution by specifying matter distribution & charge distribution. Buchdahl[2] has also considered some regular general relativistic charge fluid spheres. Some other cases of interior solutions for charged fluid sphere have been presented by Baliyn[1], Cooperstock & Cruz[3], Nduka[4], Yadav & Purushottam[8].

In this paper considering spherically symmetric metric we have obtained some solutions of Einstein-Maxwell field equation using different assumptions on metric potentials  $g_{11}$  and  $g_{44}$ . The pressure, matter, density electric field & charge density have been found.

## 2. The Field Equations

We consider the metric in the form

$$(2.1) \quad ds^2 = e^\beta dt^2 - e^\alpha dr^2 - r^2(d\theta^2 + \sin^2\theta d\phi^2)$$

Where  $\alpha$  and  $\beta$  are functions of  $r$  only.

The Einstein- Maxwell field equations for the charged perfect fluid distribution are

$$(2.2) \quad R_{ij} - \frac{1}{2}Rg_{ij} = -8\pi T_{ij}$$

$$(2.3) \quad F_j^i = 4\pi J^i = 4\pi\sigma u^i$$

$$(2.4) \quad F_{[ij;k]} = 0$$

Where  $T_{ij}$  is the energy momentum tensor,  $J^i$  is the charged current four vector,  $R_{ij}$  is the Ricci tensor and  $R$  the scalar of curvature tensor. For the system under study the energy momentum tensor  $T_j^i$  splits up to into two part viz.  $\bar{T}_j^i$  and  $E_j^i$  for matter and charges respectively.

$$(2.5) \quad T_{ij} = \bar{T}_j^i + E_\beta^\alpha$$

where

$$(2.6) \quad \bar{T}_j^i = [(\rho + p)u^i u_j - p\delta_j^i]$$

With

$$(2.7) \quad u^i u_i = 1$$

The non- vanishing component of  $\bar{T}_j^i$  are

$$(2.8) \quad \bar{T}_1^1 = \bar{T}_2^2 = \bar{T}_3^3 = -p \text{ and } \bar{T}_4^4 = \rho$$

Here  $p$  is internal pressure,  $\rho$  and  $\sigma$  are densities of matter and charges respectively,  $u^i$  is the velocity vector of the matter.

The static condition is given by

$$(2.9) \quad u^1 = u^2 = u^3 = 0 \text{ and } u^4 = (g_{44})^{-1/2}$$

i.e  $u^4 = e^{-\beta/2}$

The electromagnetic energy momentum tensor  $E_j^i$  is given by

$$(2.10) \quad E_j^i = -F_{jk} F^{ik} + \frac{1}{4} \delta_j^i F_{lm} F^{lm}$$

We assume the field to be purely electrostatic i.e  $F_{ik} = 0$  and  $F_{4k} = \phi, k = \phi_k$  and where  $\phi$  is the electrostatic potential.

Thus the Einstein-Maxwell field equations are reduced into form

$$(2.11) \quad e^{-\alpha} \left( \frac{1}{r^2} - \frac{\alpha}{r} \right) - \frac{1}{r^2} = -8\pi\rho - E$$

$$(2.12) \quad \frac{1}{r^2} - e^{-\alpha} \left( \frac{1}{r^2} + \frac{\beta'}{r} \right) = -8\pi p + E$$

$$(2.13) \quad e^{-\alpha} \left[ \frac{1}{4} \beta' \alpha' - \frac{1}{4} \beta'^2 - \frac{1}{2} \beta'' - \frac{1}{2} \left( \frac{\beta' - \alpha'}{r} \right) \right] = -8\pi p - E$$

Where

$$(2.14) \quad E = -F^{41} F_{41}$$

$$(2.15) \quad 4\pi\rho = \left( \frac{\partial F^{41}}{\partial r} + \frac{2}{r} F^{41} + \frac{\alpha' + \beta'}{2} F^{41} \right) e^{\beta/2}$$

### 3. Solution of the Filled Equations

We have four equations (2.11)-(2.13) and (2.15) in six variables  $(\rho, E, p, \alpha, \beta, \sigma)$ . Here we take  $\alpha$  and  $\beta$  are two free variables.

We choose

$$(3.1) \quad e^\alpha = \frac{Ar^3 + Br^2 + C}{r^3 + C}$$

$$(3.2) \quad e^\beta = \frac{Lr^3 + Mr + C}{4N}$$

Where A, B, C, L, M, N are arbitrary constants

Pressure, matter density, electric field and charge density are found to be

$$(3.3) \quad 8\pi p = \frac{r^3 + C}{2(Ar^3 + Br^2 + C)} \left[ \frac{36L^2r^5 + 54LM^4 + 32CLr^2 + 5M^2r^2 + 6CM}{4r(Lr^3 + M + C)^2} \right]$$

$$(3.4) \quad 8\pi\rho = \frac{r^3 + C}{2(Ar^3 + Br^2 + C)} \left[ \frac{\frac{9}{4}r^3 C(A-1) + (12L^4 + 11Mr^2 + 10Cr)}{2r(r^3 + C)(Ar^3 + Br^2 + C)(Lr^3 + Mr + C)} + \frac{4L^2r^5 + 6LMr^4 + 3M^2r + 2CMr}{4r(Lr^3 + M + C)^2} - \frac{1}{r^2} \right] + \frac{1}{2r^2}$$

$$(3.5) \quad E = \frac{r^3 + C}{2(Ar^3 + Br^2 + C)} \left[ \frac{\frac{9}{4}r^3 C(1-A)\{Lr^3 + (M+2L)r + (C+M)\}}{(r^3 + C)(Ar^3 + Br^2 + C)(Lr^3 + Mr + C)} - \frac{1}{r^2} \right] + \frac{1}{2r^2}$$

$$(3.6) \quad 4\pi\sigma = \left[ \frac{\partial F^{41}}{\partial r} + \frac{2}{r} F^{41} + \frac{1}{2} \left\{ \frac{3r^2(CA - Br^2 + C) + 3r^4B + 2CB}{(r^3 + C)(Ar^3 + Br^2 + C)} + \frac{3r^2L + M}{Lr^3 + Mr + C} \right\} F^{41} \right] X \left( \frac{Lr^3 + Mr + C}{4N} \right)^{1/2}$$

The constants appearing in the solution can be found using boundary conditions.

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